

SIDE-SLIPPING OF A RADIATING PARTICLE

X. Artru^a, G. Bignon^b, T. Qasmi^aInstitut de Physique Nucléaire de Lyon, IN2P3-CNRS Université Claude-Bernard, France

e-mail: x.artru@ipnl.in2p3.fr

^bÉcole Normale Supérieure de Lyon, France

Radiation reaction is revisited, first in a new classical approach, where the physical particle 4-momentum is redefined as the energy-momentum flux across the future light cone and is not parallel to the 4-velocity.

Then in a semi-classical approach, it is shown that, when emitting a photon, the particle "side-slips" transversally to its initial momentum, justifying the non-collinearity between momentum and mean velocity. Side-slipping is finally checked in a pure quantum mechanical treatment of synchrotron radiation.

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1. RECALL ABOUT RADIATION

An electron submitted to an external field ($\mathbf{E}_m, \mathbf{B}_m$) in vacuum emits radiation with the power:

$$\frac{dW_{rad}}{dt} = \frac{2}{3} \alpha_{cl} \left(\frac{d\mathbf{X}}{dT} \right)^2. \quad (1)$$

The relativistic 4-vector generalization is

$$\frac{dP_{rad}^\mu}{d\tau} = \frac{2}{3} \alpha_{cl} (\ddot{\mathbf{X}} \cdot \ddot{\mathbf{X}}) \dot{X}^\mu. \quad (1')$$

We take unified dimensions for space and time ($c=1$) with the $(-+++)$ metric. $\dot{\mathbf{X}} \equiv d\mathbf{X}/d\tau$, where τ is the proper time. We also use rationalized Maxwell equations, e.g. $\nabla \cdot \mathbf{E} = \rho$. We keep $\hbar \neq 1$ and define $\alpha_{cl} \equiv e^2/(4\pi\hbar) = 1/137$, where $e = -|e|$ is the charge of the electron. To account for the loss of the electron energy, Abraham and Lorentz introduced the dissipative force

$$\mathbf{f}_{reac} = \frac{2}{3} \alpha_{cl} \ddot{\mathbf{X}}. \quad (2)$$

The non-relativistic equation of motion is then

$$m \ddot{\mathbf{X}} = e (\mathbf{E}_{in} + \dot{\mathbf{X}} \times \mathbf{B}_{in}) + \mathbf{f}_{reac}$$

whose relativistic generalization is the Abraham-Lorentz-Dirac (ALD) equation:

$$m \ddot{\mathbf{X}} = e F_{in}(X) \cdot \dot{\mathbf{X}} - \frac{2}{3} \alpha_{cl} [(\ddot{\mathbf{X}} \cdot \ddot{\mathbf{X}}) \dot{\mathbf{X}} - \ddot{\mathbf{X}}] \quad (3)$$

We use the notation $(F \dot{\mathbf{X}})^\mu \equiv F^{\mu\nu} \dot{X}_\nu$. $F_{in} = \{\mathbf{E}_m, \mathbf{B}_m\}$ is the "incoming" (or "external") electromagnetic field, related to the total, retarded, advanced and outgoing fields by

$$F_{tot} = F_{in} + F_{ret} = F_{adv} + F_{out}, \quad (4)$$

$$F_{rad} = F_{out} - F_{in} = F_{ret} - F_{adv}, \quad (4')$$

In the following we shall omit the suffix *in*. An excellent review on radiation reaction can be found in Ref. [1].

The "mad electron"

Although mathematically elegant, the ALD equation is not physically acceptable for the following reasons:

* a third initial condition $\ddot{\mathbf{X}}(0)$ is needed in addition to $\mathbf{X}(0)$ and $\dot{\mathbf{X}}(0)$.

* for almost every $\ddot{\mathbf{X}}(0)$, the electron eventually goes into a *run-away* motion.

* given $\mathbf{X}(0)$ and $\dot{\mathbf{X}}(0)$, there may exist one (or a discrete set of) $\ddot{\mathbf{X}}(0)$ such that the electron avoids run-away motion, but this value depends on all the fields $F_{in}(X)$ that the electron will encounter in the future. Saying that "nature precisely chooses this $\ddot{\mathbf{X}}(0)$ " constitutes a violation of the causality principle. One may compare this situation with the following one: In a bus, a passenger puts a stick vertical on the floor and wants it to remain standing up in equilibrium during the whole journey, and also after the bus has stopped. To counteract the accelerations of the bus, he or she must give some initial angular velocity to the stick (Fig. 1). To do so, the passenger must know exactly in advance the accelerations of the vehicle during the whole journey.

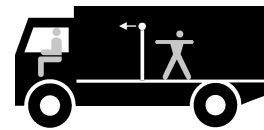


Fig. 1. Stick standing in equilibrium in a truck

The run-away instability is probably related to the point-like limit of the classical electron considered by Lorentz: For a sufficiently small radius, the electrostatic self-energy is larger than the physical mass. Then the electron "core" has a negative mass and "likes" to accelerate, since that lowers its kinetic energy.

It is possible to find approximations of the ALD equation, valid to first order in α_{cl} , which remove the arbitrariness of $\ddot{\mathbf{X}}(0)$ and have no run-away solutions. One of them [2] is obtained by replacing $\ddot{\mathbf{X}}$ and $\ddot{\mathbf{X}}$ in the right-hand side of (3) by their values calculated without radiation reaction, $\ddot{\mathbf{X}} \rightarrow (e/m)F\dot{\mathbf{X}}$,

$$\ddot{\mathbf{X}} \rightarrow (e/m)F\dot{\mathbf{X}} + (e/m)^2 FF\dot{\mathbf{X}},$$

with $\dot{F} = \dot{X}^\lambda \partial_\lambda F(X)$. One obtains

$$m\ddot{X} = eF\dot{X} + \sigma_{Th} \left[FF\dot{X} - (F\dot{X})^2 \dot{X} \right] + \frac{2e}{3} r_{cl} \ddot{F}\dot{X}, \quad (5)$$

where $r_{cl} = e^2/(4\pi m)$ is the classical electron radius and $\sigma_{Th} = (8\pi/3) r_{cl}^2$ – the Thomson cross section. The second term of (5) can be interpreted as the radiation pressure of the incoming field.

2. REFORMULATION OF THE ALD EQUATION AND NEW APPROXIMATIONS

Usually, one identifies $m\ddot{X}$ with the physical 4-momentum of the particle. Then Eqs. (1') and (3) do not seem to conserve the total 4-momentum instantaneously, because of the *Schott term* $(2\alpha_{cl}/3)\ddot{X}$. Redefining the 4-momentum as

$$P^\mu = m\dot{X}^\mu - \frac{2}{3}\alpha_{cl}\ddot{X}^\mu, \quad (6)$$

the ALD equation can be replaced by the following system:

$$m\dot{X} = P + \frac{2}{3}\alpha_{cl}\ddot{X} \quad (7a)$$

$$\dot{P} = eF\dot{X} - \frac{2}{3}\alpha_{cl}(\ddot{X} \cdot \ddot{X})\dot{X}. \quad (7b)$$

Eqs.(1') and (7b) make the instantaneous conservation of the total 4-momentum manifest. On the other hand, the mass is not conserved: $P \cdot P \neq -m^2$ and \mathbf{P} is not collinear to the electron velocity. These two features are not physically damning. (6) can be approximated by $P(\tau) \cong m\dot{X}(\tau - 2r_e/3)$, telling that the electro-magnetic part of P^μ follows the variations of the core velocity with some delay.

In what follows, we shall show that (6) is a quite natural definition of P^μ . As usual, we separate P^μ in core and electromagnetic contributions:

$$P = m_c \dot{X} + \delta P \quad (8)$$

with

$$\delta P^\mu = \int_\Sigma d\Sigma_\nu \theta^{\mu\nu} \quad (9)$$

where $\theta^{\mu\nu}$ is the energy-momentum flux tensor of the electron field. The latter field is not uniquely defined: it

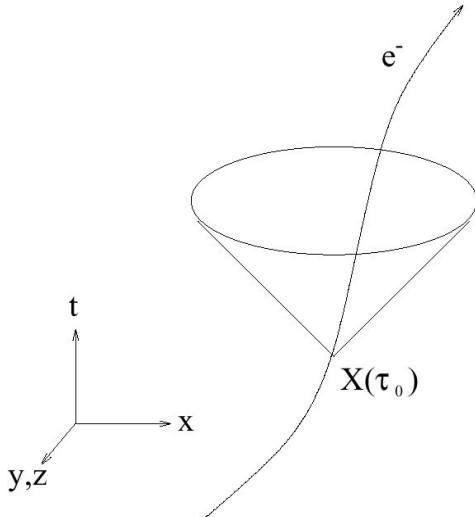


Fig. 2. World line and future light cone of the particle

can be the retarded one, the advanced one or any linear combination of the two. Looking at the first decomposition of Eq. (4), we choose the *retarded* field. So we consider that the incoming field does not contribute to the self 4-momentum and only exerts a force on the core according to the first term of (3). As hyper-surface Σ we a priori choose the future light cone of the electron (Fig. 2). This avoids a contribution from the *radiated* field F_{rad} , the 4-momentum of which flows parallel to the cone and does not cross it.

Let us consider a point $X^\mu(\tau_0)$ of the electron world line. For a space-time point Y^μ of its future light cone, we define

$$Y - X(\tau_0) \equiv R = (r, \mathbf{r}) = r(1, \hat{\mathbf{r}}).$$

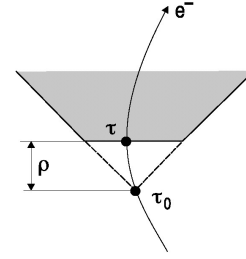


Fig. 3. Truncated lightcone or "flower-pot"

The integrand of (9) can be evaluated most easily in the electron rest frame, using standard formulas for \mathbf{E}_{ret} , \mathbf{B}_{ret} , $\Theta^{\mu\nu}$ and

$$(d\Sigma_\nu) = d^3\mathbf{r}(1, -\hat{\mathbf{r}}). \quad (10)$$

We only give the result:

$$\begin{aligned} \delta P &= \frac{\alpha_{cl}}{8\pi} \int \frac{d^3\mathbf{r}}{r^4} (1, \hat{\mathbf{r}}) = \\ &= \frac{\alpha_{cl}}{8\pi} \int \frac{d^3\mathbf{r}}{r} [\dot{X}(\tau_0) \cdot R]^{-4} R. \end{aligned} \quad (11)$$

The second expression is Lorentz invariant and applies as well in frames where the electron is not at rest. Note that the acceleration does not enter this formula. It confirms that there is no contribution of the radiated field.

The integral diverges at $r=0$, recalling that the classical self-energy of a point-like charge is infinite. In the following, we will assume that the electron has some very small but finite extension r_c . To treat the divergence, we truncate the light cone by a hyperplane orthogonal to the electron world line at $X(\tau)$ where $\tau - \tau_0 = \rho$ is a small distance, but larger than r_c (Fig. 3).

We now take the rest frame of the electron at $X(\tau)$ (not τ_0). We close the cone, truncated at $r > \rho$, by the piece of hyperplane $R^0 = \rho$, $|\mathbf{R}| \leq \rho$ and integrate (9) on the new hypersurface (in grey on Fig. 3) which we call a "flower-pot". To first order in ρ ,

$$\begin{aligned} \dot{X}(\tau_0) &= (1, -\rho \ddot{\mathbf{X}}), \\ [\dot{X}(\tau_0) \cdot r]^{-4} &= r^{-4} [1 - 4\rho \ddot{\mathbf{X}} \cdot \hat{\mathbf{r}}]. \end{aligned} \quad (12)$$

The truncated integral of (11) is

$$\begin{aligned}\delta P_{r>\rho} &= \alpha_{cl} \left(\frac{1}{2\rho}, -\frac{2}{3} \ddot{\mathbf{X}}(\tau) \right) = \\ &= \frac{\alpha_{cl}}{2\rho} \dot{\mathbf{X}}(\tau) - \frac{2}{3} \alpha_{cl} \ddot{\mathbf{X}}(\tau),\end{aligned}\quad (13)$$

the second expression being frame-independent. The hyperplane piece (bottom of the flower-pot) is the interior of the sphere of radius ρ at fixed time. It is approximately centered at $\mathbf{X}(\tau)$, the displacement being of second order in ρ . Its contribution to (9) is

$$\delta P_{r<\rho} = (\delta m_{r<\rho}, \mathbf{0}) = \delta m_{r<\rho} \dot{\mathbf{X}}(\tau). \quad (14)$$

The total 4-momentum at proper time τ (not τ_0) is obtained from (8), (13) and (14). We recover the new definition

$$P(\tau) = m\dot{\mathbf{X}}(\tau) - \frac{2}{3} \alpha_{cl} \ddot{\mathbf{X}}(\tau) \quad (15)$$

where

$$m = m_c + \delta m_{r<\rho} + \frac{\alpha_{cl}}{2\rho}$$

is the renormalized mass of the electron. The third term is the Coulomb self-energy at $r \geq \rho$ whereas the detailed short range structure of the electron is summarized in the sum of the first two terms.

Let us make the energy-momentum balance in the space-time region between two successive "flower-pots" at proper times τ and $\tau + d\tau$:

- * $P(\tau)$ is coming through the first flower-pot,
- * $dP_{in} = e F_{in} \dot{\mathbf{X}} d\tau$ is brought to the core by F_{in} ,
- * $dP_{rad} = \frac{2}{3} \alpha_{cl} (\ddot{\mathbf{X}} \cdot \ddot{\mathbf{X}}) \dot{\mathbf{X}} d\tau$ is radiated at infinity

between the two flower-pots,

$$* P(\tau + d\tau) = P(\tau) + \dot{P} d\tau$$

is outgoing through the second flower-pot.

Adding the first two quantities and subtracting the last two ones must give zero. This gives (7b).

The above calculations constitute a new and relatively simple derivation of the ALD equation, written in the form (7). From this form on can derive new types of approximations [3-5], also valid to first order in α_{cl} . The simplest one to implement in a computer code is obtained replacing $\ddot{\mathbf{X}}$ in the right-hand sides by $(e/m^2)FP$. One may in addition replace the last $\dot{\mathbf{X}}$ by P/m . Compared to (5), these approximations have the advantage of not involving the field derivatives.

3. SEMI-CLASSICAL APPROACH

Eq. (6) tells that the momentum does not follow the velocity, but one may see things the other way around and say that the electron does not follow the direction of its momentum. We call this phenomenon side-slipping, by analogy with a skier whose track is not always tangential to the skis (Fig. 4)



Fig. 4. Side-slipping skier

A discrete side-slipping is naturally obtained in a semi-classical description of the process $e^- \rightarrow e'^- + \text{photon}$ in an external field. If we consider this process as instantaneous and local at a definite point \mathbf{X} of the trajectory, it cannot satisfy the conservation of both momentum,

$$\mathbf{P} = \mathbf{P}' + \mathbf{k} \quad (16)$$

and energy

$$\mathbb{P} = \mathbb{P}' + \mathbb{P}_\omega \quad (17)$$

with $\mathbb{P} \equiv \sqrt{\mathbf{P}^2 + m^2}$. However, 4-momentum conservation becomes possible if we assume that the final electron trajectory starts from a point $\mathbf{X}' \neq \mathbf{X}$ aside from the initial trajectory. In the case of a static electric field, we replace (17) by

$$U(\mathbf{X}') - \mathbb{P}' = U(\mathbf{X}) - \mathbb{P} + \mathbb{P}_\omega \quad (17')$$

where $U(\mathbf{X})$ is the potential energy. (17') and (17) give

$$\delta U = U(\mathbf{X}') - U(\mathbf{X}) \cong \frac{\mathbb{P}}{\Gamma} \mathbb{P}_\omega \frac{\gamma^{-2} + \theta^2}{2} \quad (18)$$

where θ is the angle between \mathbf{P} and \mathbf{k} . δU is obtained by a finite displacement $\delta \mathbf{X} = \mathbf{X}' - \mathbf{X}$ of the electron toward a lower potential energy. In the ultrarelativistic case, we take $\delta \mathbf{X}$ perpendicular to the trajectory:

$$\delta \mathbf{X} = - \frac{\mathbb{P}}{\Gamma} \mathbb{P}_\omega \frac{\gamma^{-2} + \theta^2}{2} \frac{\mathbf{f}_\perp}{|\mathbf{f}_\perp|^2}, \quad (19)$$

where \mathbf{f}_\perp is the transverse component of the force. Such a "side-slipping" was already introduced in channeling radiation (Eqs. (15)-(17) of Ref. [6]). It contributes to the decrease of the transverse energy which explains the very fast energy loss of axially channeled electrons above hundred GeV.

Let us now consider synchrotron radiation in a

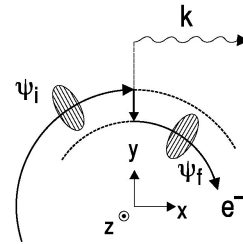


Fig. 5. Photon emission in a synchrotron

uniform magnetic field $\mathbf{B} = -B\hat{z}$ derived from the vector potential

$$\mathbf{A} = (yB, 0, 0). \quad (20)$$

The particle hamiltonian is $(\mathbf{P}^2 + m^2)^{1/2}$ where $\mathbf{P} = \mathbf{p} - e\mathbf{A}$ is the mechanical momentum and \mathbf{p} the canonical one. In the gauge (20), the hamiltonian is invariant under translation in the x and z directions, therefore p_x and p_z are conserved. We assume that the photon is emitted when the electron is at $x=0, y=R$ (Fig. 5). Then we require the conservation laws (16-17),

but with \mathbf{p} and \mathbf{p}' in place of \mathbf{P} and \mathbf{P}' . For the x -component it writes

$$P_x + eyB = P'_x + ey'B + \square k_x, \quad (21)$$

where we have anticipated a side-slipping $y' = y + \delta y$. For $|e|B\delta y$ we obtain the same result (18) as for δU and, since $|e|B > |\mathbf{f}|$, δy is given by (19) or

$$\delta y = \frac{\square \omega}{\Gamma} R \frac{\gamma^{-2} + \theta^2}{2}. \quad (22)$$

The side-slipping has also the virtue of insuring angular momentum conservation. Let us consider again the circular trajectory of Fig. 5, but now due to the spherically symmetric potential $U(|\mathbf{X}|)$. Neglecting spin, the z -component of the angular momenta of the initial and final electrons are

$$L_z = -yP_x, \quad L'_z = -y'P'_x. \quad (23)$$

Here we neglect the quantum recoil effect, i.e. we use the classical or *soft photon* approximation ($\hbar\omega \ll m$). The source of the radiation – and the radiation itself – is invariant under a time translation by $\square t$ times a rotation by the angle $v\square t/R$. For a photon quantum state of definite angular momentum J_z and frequency ω , this invariance is expressed as

$$\exp[-i(v\square t/R)J_z] \times \exp(-i\omega \square t) = 1$$

therefore

$$J_z = -\omega R/v. \quad (24)$$

The conservation of angular momentum,

$$L_z = L'_z + J_z. \quad (25)$$

together with that of linear momentum along $\hat{\mathbf{x}}$,

$$P_x = P'_x + \square k_x \quad (26)$$

yield the result (22) again, with $\square > \square'$.

Incidentally, identifying (24) with the "classical photon" result $J_z = -yk_x$ implies a "side-slipping" for the photon also:

$$y_{phot} - R = \frac{\gamma^{-2} + \theta^2}{2} R, \quad (27)$$

which could be observed at low-energy synchrotron machines.

The side-slipping formula (19) can be generalized in a covariant form, writing the 4-momentum conservation as

$$P + Q = P' + \square K \quad (28)$$

We assume that Q is provided by the work of the external field along δX :

$$Q = eF\delta X. \quad (29)$$

Squaring the two sides of (28), using $P^2 = P'^2 = -m^2$, $K^2 = 0$ and neglecting Q^2 a priori, we obtain

$$P \cdot Q = \square K \cdot P' \quad \left(\cong \frac{\square}{\Gamma} \square K \cdot P \right). \quad (30)$$

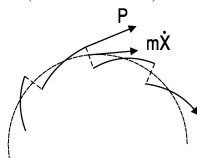


Fig. 6. Semi-classical electron trajectory emitting photons successively

Using $P \cdot F \dot{P} = -\dot{P} \cdot F P$ and, to first order in α_{cl} ,

$P = m\dot{X}$, $F\dot{X} = \dot{P}$, one can verify that

$$\delta X^\mu = \frac{\square - K \cdot P'}{m \cdot P \cdot P} \dot{P}^\mu \quad (31)$$

inserted in (29), satisfies (30). The neglect of Q^2 has to be checked a posteriori from (29). We expect it to be small if the external field varies smoothly, e.g. in synchrotron or channeling radiations, interpreting Q as the momentum of the virtual photon(s) taken from the external field.

In the limit $\hbar \rightarrow 0$, the 4-momenta of the individual photons goes to zero and their number goes to infinity so that the total radiated 4-momentum is finite and given by (1'). Summing all the small side-slippings (31) during the proper time dt , approximating $K \cdot P'$ by $K \cdot P$, one recovers Eq. (7a), to first order in α_{cl} . This is illustrated in Fig. 6.

4. FULL QUANTUM DERIVATION

Side-slipping was deduced above from semi-classical arguments of energy, momentum and angular momentum conservation. Here we will derive side-slipping from a full quantum treatment, in the particular case of synchrotron radiation. Neglecting electron spin, we start from the Klein-Gordon equation (now $\hbar = 1$),

$$[(\nabla - ie\mathbf{A})^2 - \partial_t^2 - m^2]\psi = 0 \quad (32)$$

and consider a wave packet of the form

$$\psi = e^{ipx - i\square t} \psi \quad (33)$$

where $P^\mu = (\square, p, 0, 0)$ is a reference 4-momentum and ψ a slowly varying function of (t, x, y, z) . Using $\square^2 = p^2 + m^2$, (32) becomes

$$\left[i\square \partial_t - \partial_t \partial_x + i(E - p)\partial_x + \partial_y^2 + \partial_z^2 + 2fy(p - i\partial_x) - f^2 y^2 \right] \psi = 0 \quad (34)$$

where $\partial_\pm = \partial_t \pm \partial_x$ and $f = |e|B$. Assuming $\square > p \square m$ we consider ∂_+ to be of order \square^{-1} , which allows us to neglect the second and third terms of the square bracket. Furthermore, we take a wave packet located near $(x, y, z) = 0$ at time $t = 0$ (we change the origin of the coordinates in Fig. 5). So we neglect the terms in y^2 and $y\partial_x$ (but not in yp). We get

$$\left[i(\partial_t + \partial_x) + \frac{1}{2\square} (\partial_y^2 + \partial_z^2) - fy \right] \psi = 0. \quad (35)$$

Looking for solution of the form

$$\psi(t, x, y, z) = \chi(x - t)\phi(t, y, z), \quad (36)$$

we are left with the 2-dimensional Schrödinger equation for a particle of mass $\square = \gamma m$ in the linear potential $V(y) = by$:

$$\left[i\partial_t + \frac{1}{2\square} (\partial_y^2 + \partial_z^2) - fy \right] \phi = 0. \quad (37)$$

Using the coordinate of the accelerated frame

$$y_a = y + \frac{f}{2\square} t^2 \quad (38)$$

and setting

$$\phi(t, y, z) = \phi_a(t, y_a, z) \exp\left(-ifty - i \frac{b^2 t^3}{6\Omega}\right) \quad (39)$$

we transform (37) in the free-particle Schrödinger equation,

$$\left[i\partial_t + \frac{1}{2\Omega} \left(\partial_{y_a}^2 + \partial_z^2 \right) \right] \phi_a(t, y_a, z) = 0. \quad (40)$$

Thus ϕ_a can be expanded in plane waves:

$$\phi_a = \iint \frac{dq}{2\pi} \frac{dr}{2\pi} \tilde{\phi}_a(q, r) \exp\left(iqy_a + irz - i \frac{q^2 + r^2}{2\Omega} t \right) \quad (41)$$

To sum up,

$$\begin{aligned} \Psi &= e^{ipx - i\Omega t} \chi(x - t) \exp\left(-ifty - i \frac{b^2 t^3}{6E}\right) \times \\ &\times \iint \frac{dq}{2\pi} \frac{dr}{2\pi} \tilde{\phi}_a(q, r) \exp\left[i \left(qy + \frac{qft^2}{2\Omega} + rz - \frac{q^2 + r^2}{2\Omega} t \right) \right]. \end{aligned} \quad (42)$$

We consider the transition from the electron state $|i\rangle = |\Psi\rangle$ (in the Schrödinger representation) to the electron + photon state $|f\rangle = |\Psi'\rangle \otimes |\mathbf{k}, \mathbf{a}\rangle$ where Ψ' is given by (42) with primed quantities. The wave packets Ψ and Ψ' are represented by the striated ellipses of Fig. 5. Taking (without loss of generality) \mathbf{k} along the x axis, the photon vector potential is given by

$$\mathbf{A}(t, \mathbf{X}) = \mathbf{a} e^{ik(x-t)}. \quad (43)$$

To first order in perturbation, the transition amplitude is

$$\langle f | S | i \rangle = -i \int dt \langle f | H_I | i \rangle \quad (44)$$

with the interaction hamiltonian H_I given by

$$\langle f | H_I | i \rangle = ie \int d^3 \mathbf{X} \Psi'^*(t, \mathbf{X}) \mathbf{A}^*(t, \mathbf{X}) \cdot \nabla \Psi(t, \mathbf{X}). \quad (45)$$

We now combine Eqs.(42-45). Integrations over y and z impose $q = q'$ and $r = r'$. Using the shifted variable $x' = x - t$, the integration over the x -dependent factors gives

$$e^{i(p'+k-p)t} \int dx' e^{i(p'+k-p)t} \chi(x') \chi'^*(x'). \quad (46)$$

We introduce the parameter

$$\Lambda = (\Omega' - p' + p) / m^2 \cong \frac{k}{2\Omega\Omega'} \quad (47)$$

and write the remaining 3-fold integral as

$$\begin{aligned} I &= \iiint dt dq dr \tilde{\phi}_a^*(q, r) \left(q \cdot \mathbf{a}_y^* + r \cdot \mathbf{a}_z^* \right) \tilde{\phi}_a(q, r) \times \\ &\times \exp\left[i\Lambda \left[(m^2 + q^2 + r^2) t - qft^2 + f^2 \frac{t^3}{3} \right] \right], \end{aligned} \quad (48)$$

Shifting the time variable $t' = t - q/f$, we can decouple the exponential into

$$\exp\left\{ i\Lambda \left[(m^2 + r^2) t' + f^2 \frac{t'^3}{3} \right] \right\} \quad (49)$$

and

$$\exp\left\{ i \frac{\Lambda}{f} \left[(m^2 + r^2) q + \frac{q^3}{3} \right] \right\}. \quad (50)$$

The phase factor of (49) is the same as in the semi-classical radiation formula,

$$\exp\left\{ i \frac{\Omega}{\Omega'} [\omega t' - \mathbf{k} \cdot \mathbf{X}(t')] \right\}, \quad (51)$$

knowing that the transverse components of the velocity $d\mathbf{X}/dt'$ are $v_y(t') = -ft'/\Omega$, $v_z = r/\Omega$. The factor Ω/Ω' is a recoil correction.

The factor which interests us is (50). Linearizing the cubic term about the mean value $\langle q \rangle = \Omega v_y$ and replacing r by $\langle r \rangle = \Omega v_z$, we can rewrite (50) as

$$C \cdot \exp(-iq\delta y) \quad (52)$$

where δy is equal to the right-hand side of (19) or (22). In (52) we recognize the operator of the y -translation by δy , written in the momentum space representation. The maximum transition amplitude is obtained when the wave packet $\tilde{\phi}'_a$ is transversally shifted from $\tilde{\phi}_a$ by δy . This confirms the semi-classical derivation of the side-slipping.

5. CONCLUSION

In this study, we have got new insight in the radiation mechanism. Using purely classical, semi-classical and quantum-mechanical approaches, we have shown that the velocity and the properly defined momentum of the radiating particle are not parallel, as illustrated in Fig. 6.

The classical run-away problem still remains unsolved, but we have obtained a new approximation of the ALD equation, without run-away and not involving the field derivatives. It can be easily implemented on a computer code.

The discrete side-slipping accompanying the emission of a photon is of the order of the compton wavelength, hence hardly detectable. However its contribution to the decrease of the transverse energy of a high-energy electron channeled in crystals may be non-negligible. The "side-slipping of the photon" (27), much larger than the electron one, may be observed with precise optics.

The transverse jumping of the particle from the initial to the final trajectory has no classical counterpart. It can be viewed as a tunnel effect. A similar effect should take place in the crossed reaction $\gamma \rightarrow e^+ + e^-$ in a strong field (Eq. 2 of Ref. [7]).

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